

# Unifying the Universe

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**PITP public lecture, UBC,**  
*Vancouver, 12 October, 2016*

The efforts to understand the universe is one of the very few things  
that lifts human life a little above the level of farce...

S. Weinberg, 1977



I don't know why anyone would  
want to study the expansion of the  
universe.

— *Ronald Reagan* —

AZ QUOTES

# Scales in the Universe

## Size of the Universe

90 billion light years »  **$10^{28}$**  cm

## Cosmology, astrophysics

**1000000000000000000000000000000 cm**

# Galaxies and their clusters

$10\text{-}100 \text{ Mpc} \gg 10^{25} \text{ cm}$

## Stars, orbit of Earth

150 million km »  **$10^{13}$**  cm

[illegible]

## Chemistry, atomic and nuclear physics

Molecules

1 angstrom »  $10^{-8}$  cm

# Atomic nuclei

1 fermi »  $10^{-13}$  cm

## High energy particle physics

## Higgs boson

»  $10^{-16}$  cm

[illegible]

# Quantum gravity

## Planck length

»  $10^{-33}$  cm

[illegible]

## Constants of Nature and Planck scale

**C' 300 000 km/sec**

speed of light



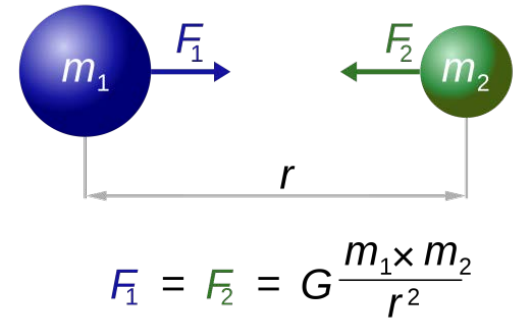
## Planck constant

$$E = \frac{\hbar c}{\lambda}$$

photon wavelength  
(color of light)



gravitational  
(Newton) constant



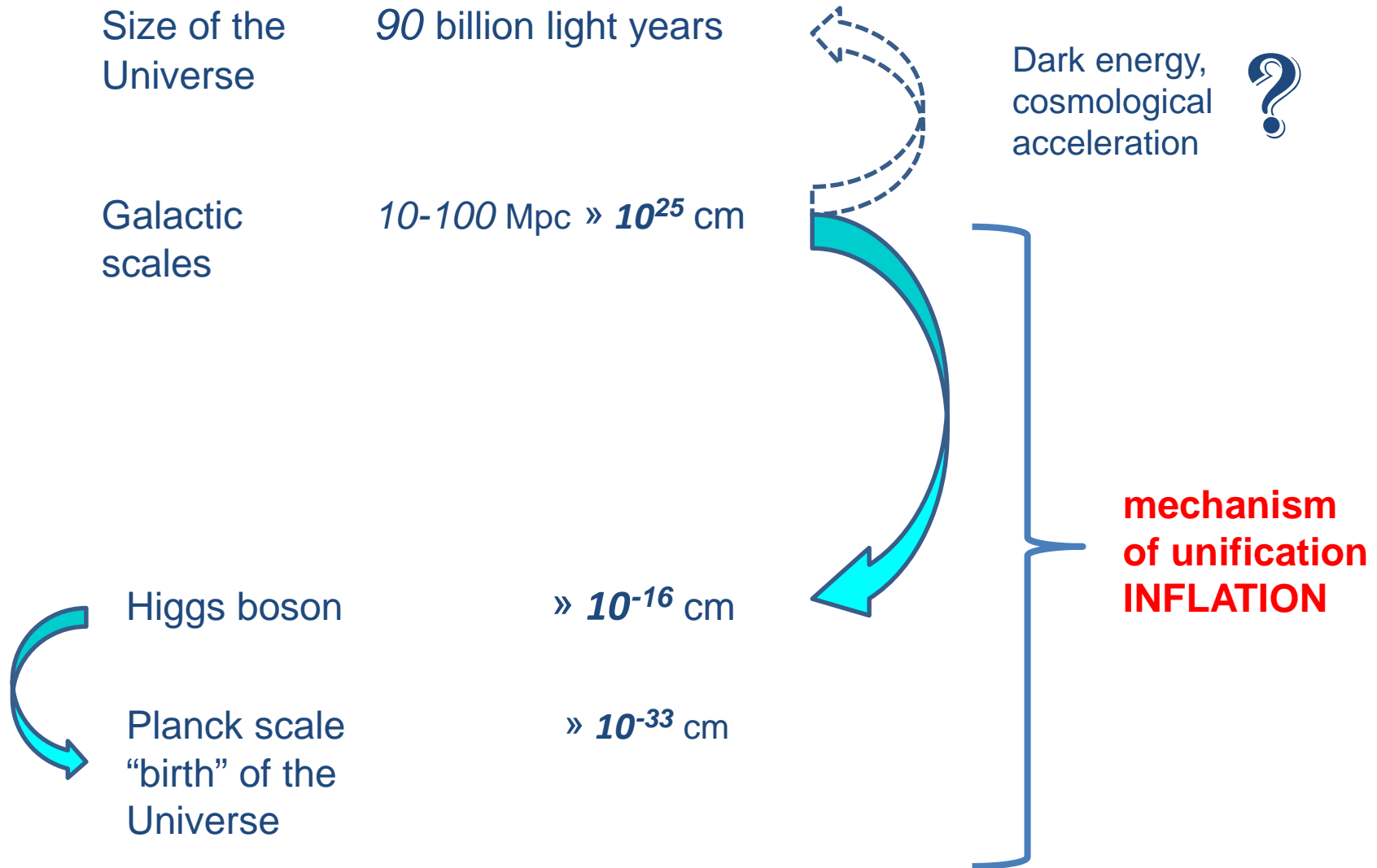
$$F_1 = F_2 = G \frac{m_1 \times m_2}{r^2}$$

**Planck length**  $l_P = \sqrt{\frac{G\hbar}{c^3}}$

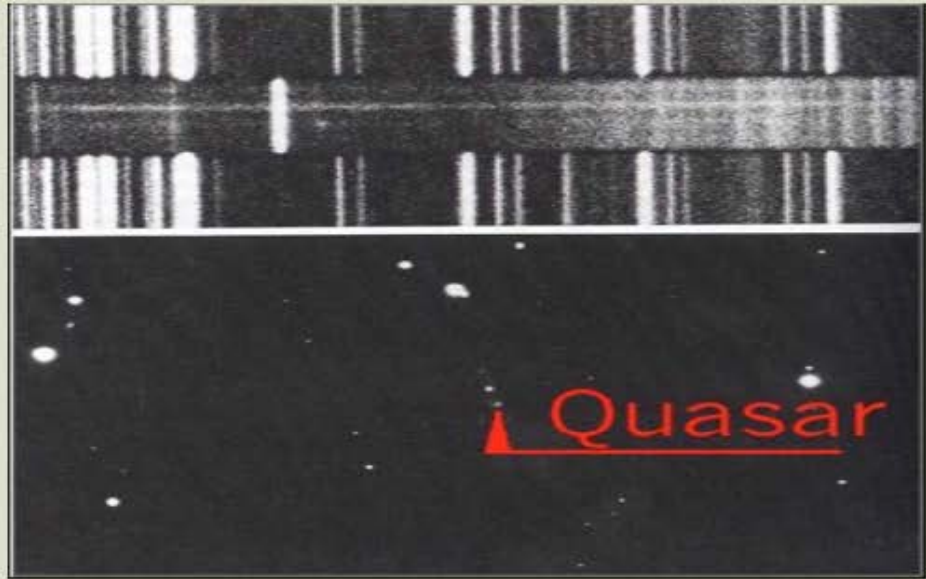
**Planck energy**  $E_P = \frac{\hbar c}{l_P} \sim 10^{19} GeV \sim 10^{19}$  protons

**100000000000000000000** Hydrogen atoms  
energy enough to send 100 kg body into space

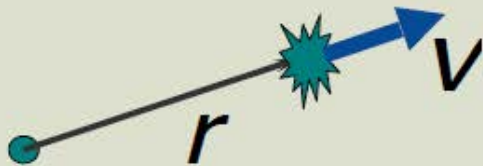
# Unification of scales:



# ● The Universe expands



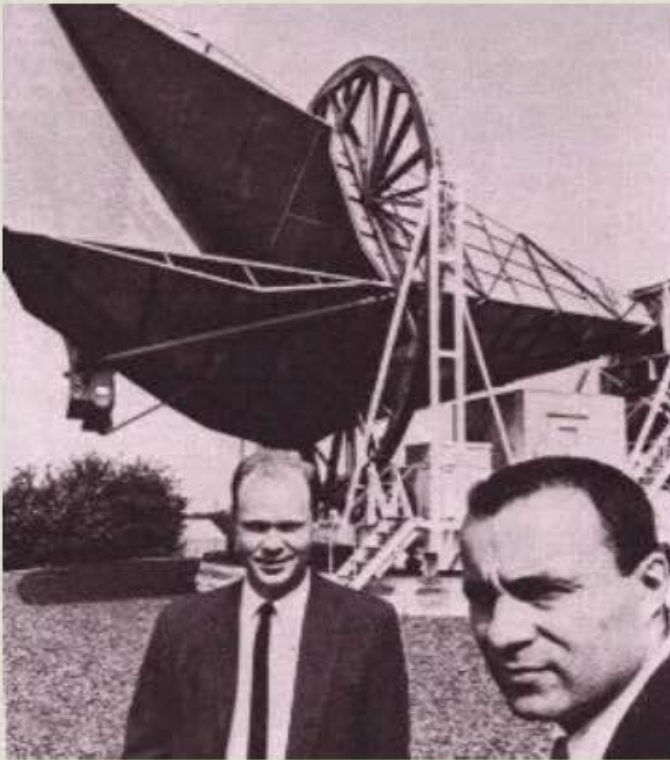
## ● Hubble law



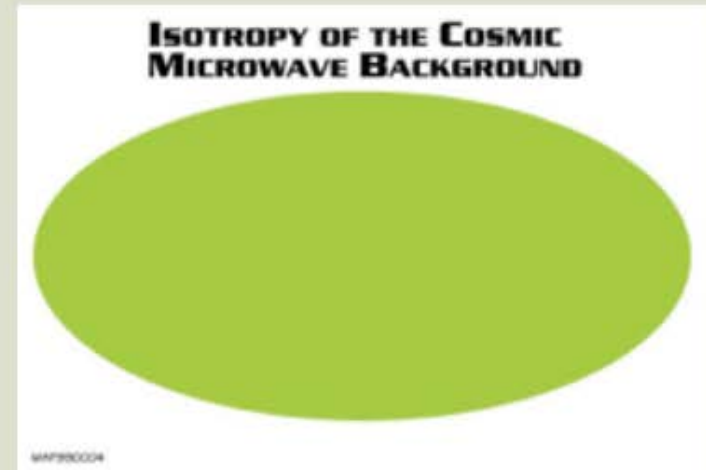
$$v = Hr$$

$$t \sim \frac{r}{v} = \frac{1}{H} \sim 13,7 \text{ bil. years}$$

- There exists background radiation with the temperature  $T \approx 3K$

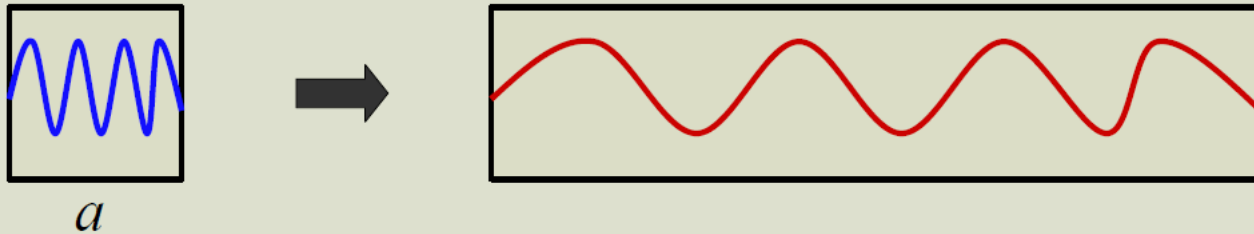


Penzias, Wilson 1965



$$\frac{\Delta T}{T} < 10^{-5} = \frac{1}{100000}$$

## Decrease of temperature (of radiation – photon gas) during expansion

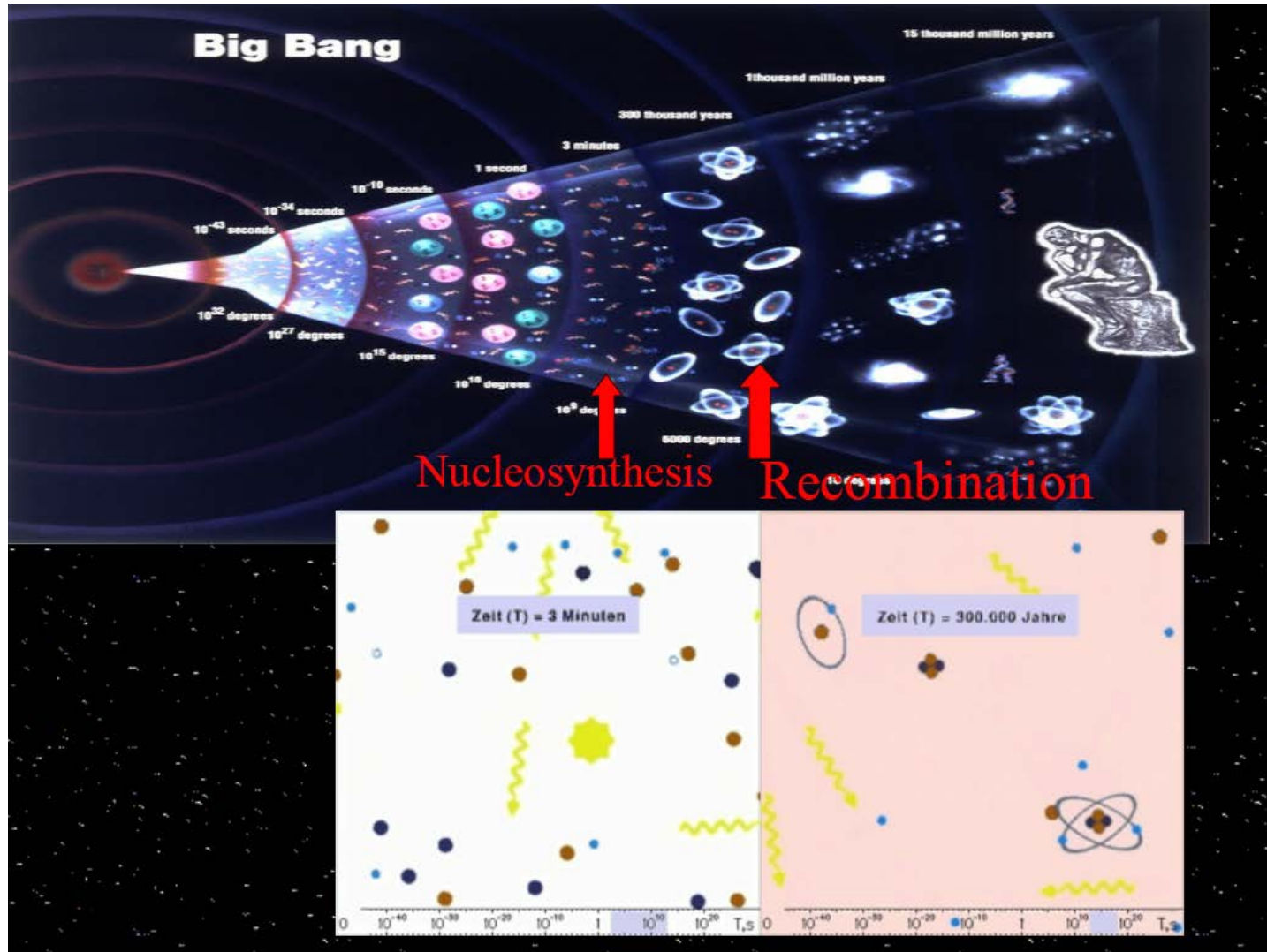


$$\lambda \propto a \quad \longrightarrow \quad T \propto \frac{1}{a}$$

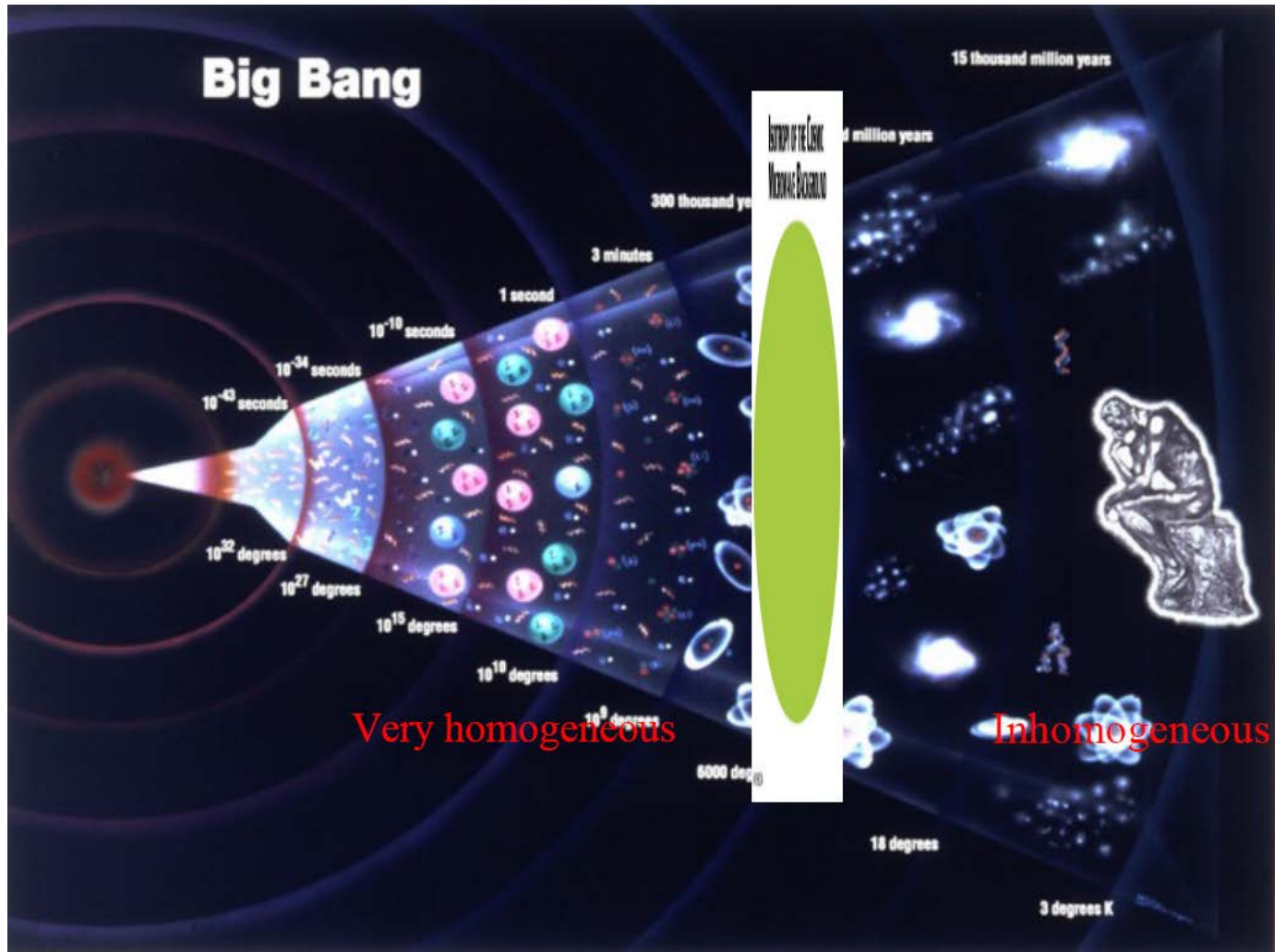
When the Universe was 1000 times smaller  
its temperature was about  $2725^{\circ}K$



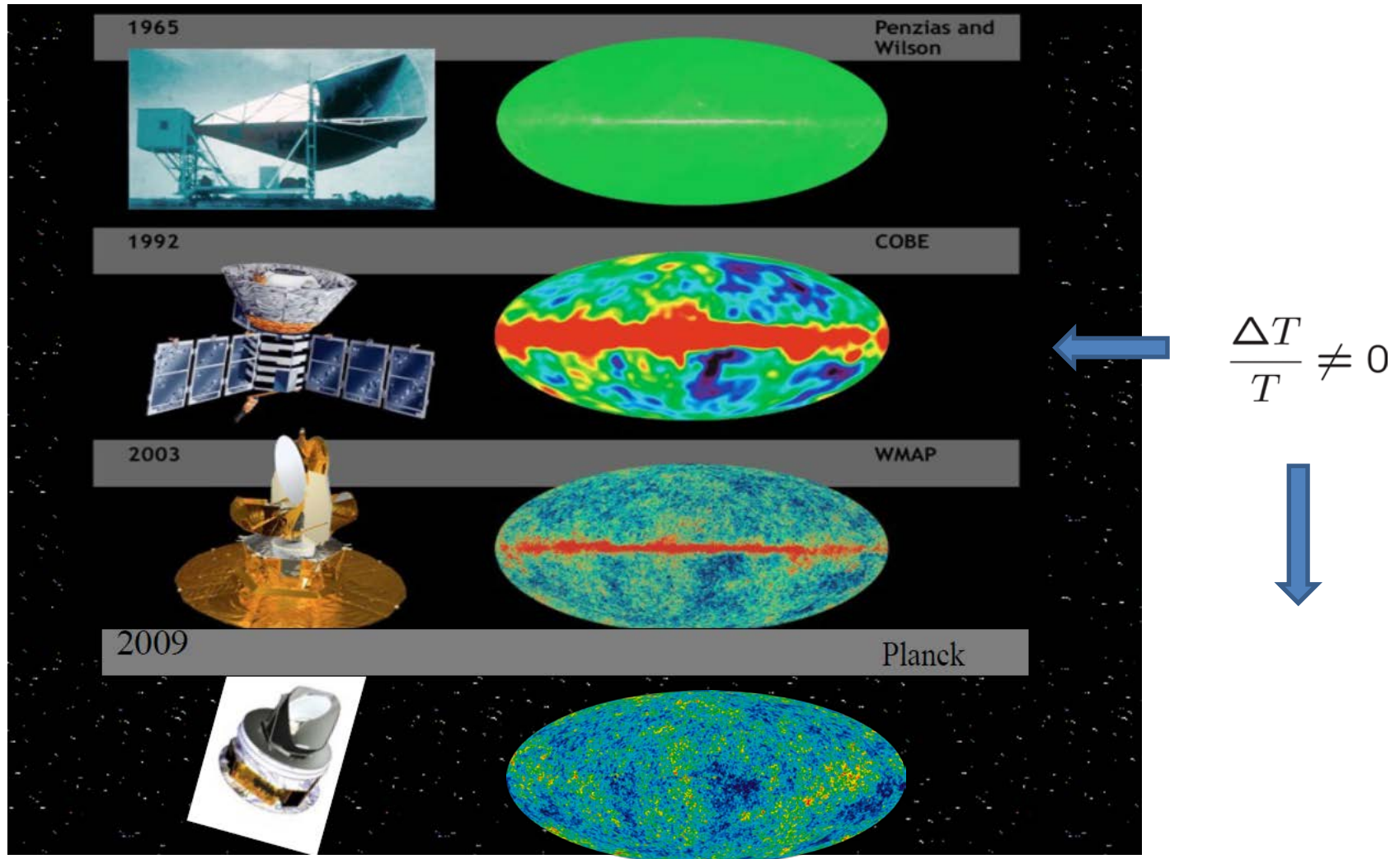
# The schematic picture of Big Bang evolution of the Universe



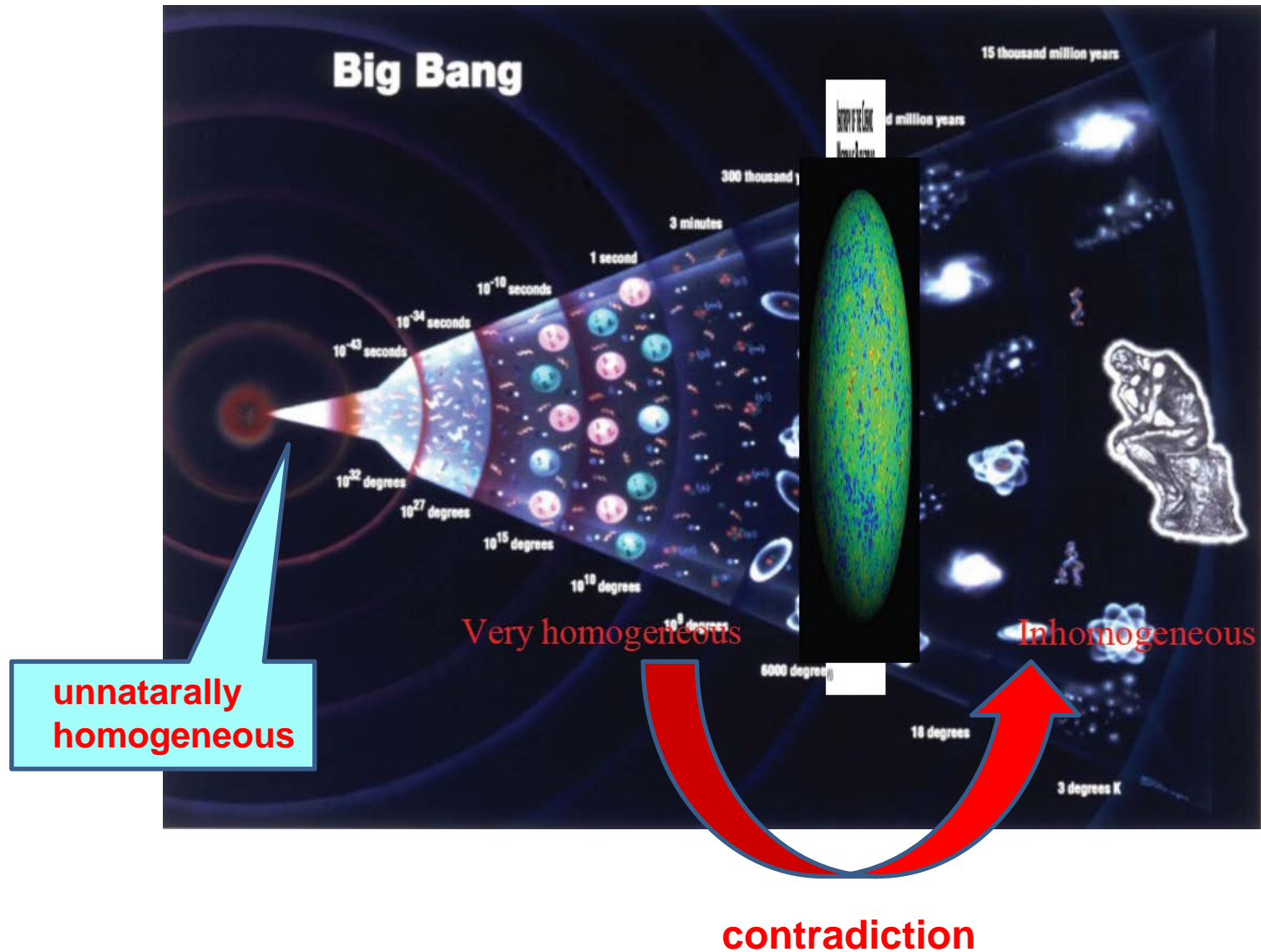
# Formation of the background (relic) radiation at the epoch of recombination



# CMB observations: 1990-ies and on – the epoch of “precision” cosmology

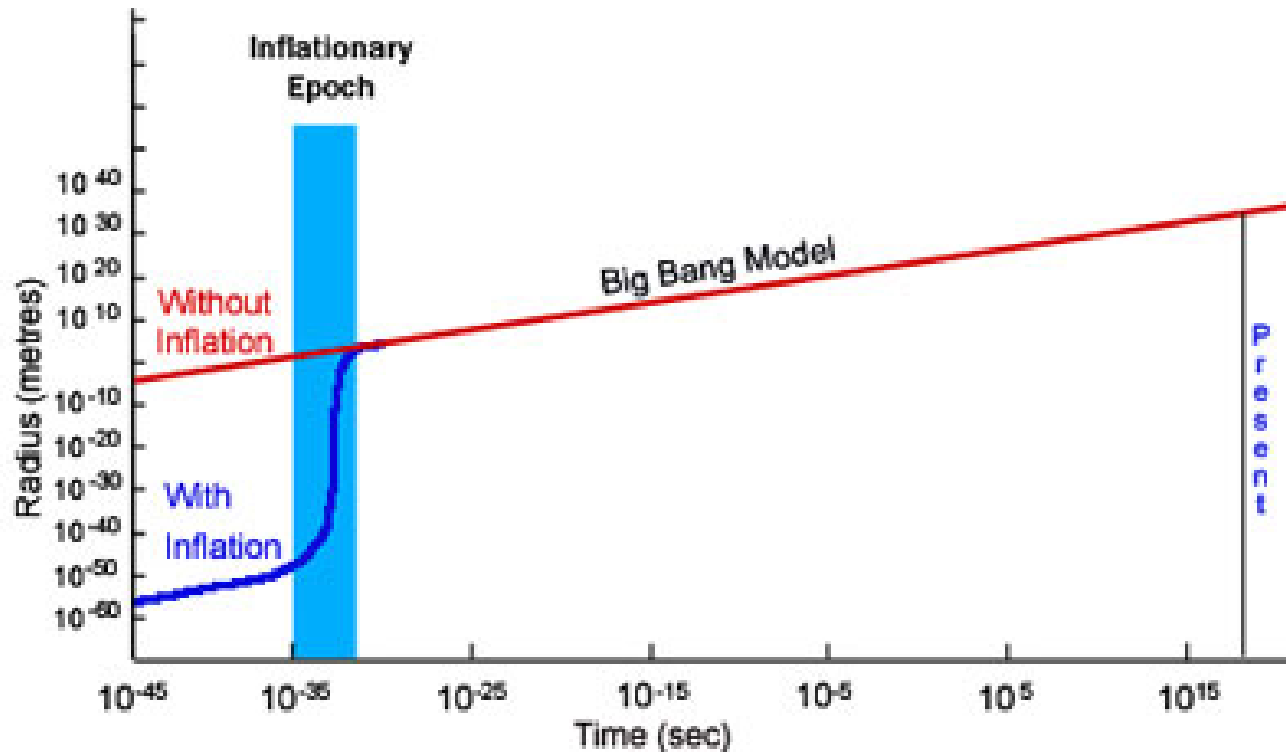


# Problems with the **old** Big Bang scenario





# Stage of rapid **accelerated** expansion – **inflation** (A.Guth)



$$a \sim e^{Ht}$$

*Exponential expansion law: when time grows twice the size gets squared, when time triples – the size gets raised to cubic power, etc.*

$$H \sim \text{const}$$

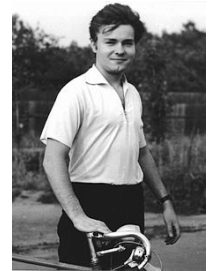
*Hubble parameter of the inflation  $\neq$  present Hubble*

# Theory of cosmological perturbations in the exponentially expanding Universe

A.Starobinsky, 1978, *gravitational waves*

V.Mukhanov and G.Chibisov, 1981,  
*waves of density perturbations as seeds  
that grew to become galaxies -- the source  
for a formation of LSS (large scale structure  
of the Universe)*

Hawking, Guth, Pi, Steinhardt, Turner,  
Albrecht



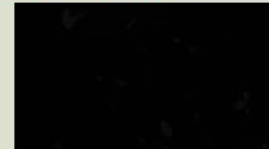
*The theory is based on the  
Heisenberg **uncertainty**  
**relation** of quantum mechanics*

*Change of paradigm: initial state is  
a **vacuum** with only quantum  
fluctuations – everything originates  
from them !*



$$\rightarrow \Delta p \Delta x \geq h$$

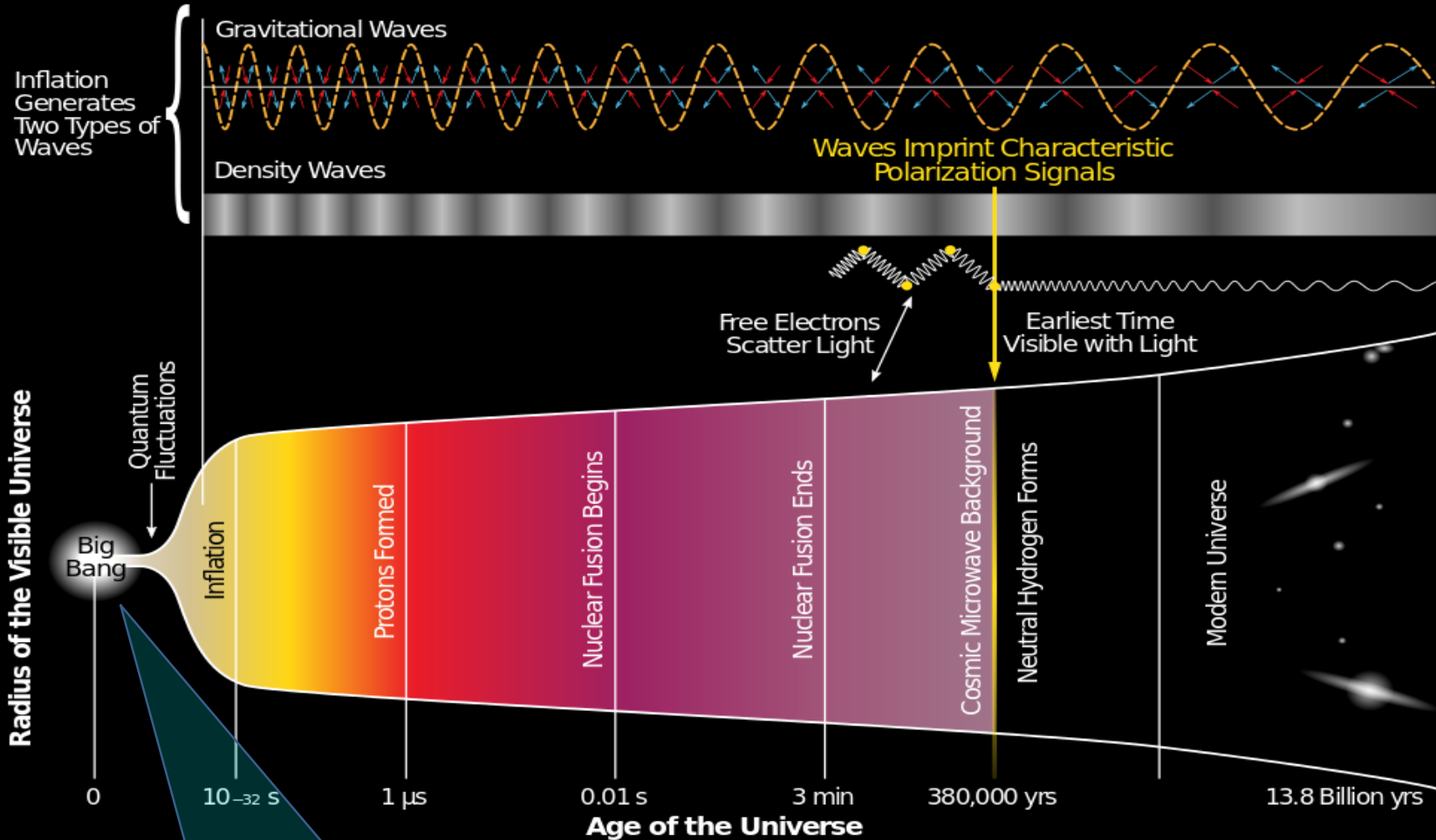
There always exist **unavoidable**  
Quantum Fluctuations



Quantum fluctuations in the density distribution are large ( $10^{-5}$ )  
only in extremely small scales ( $\sim 10^{-33}$  cm),  
but very small ( $\sim 10^{-58}$ ) on galactic scales ( $\sim 10^{25}$  cm)

Can we transfer the large fluctuations from extremely  
small scales to large scales???

# History of the Universe



a vacuum or nearly vacuum  
state of zero temperature

## Perturbations of matter and gravitational field:

*produce perturbations of temperature of relic radiation – CMB – which we observe now*

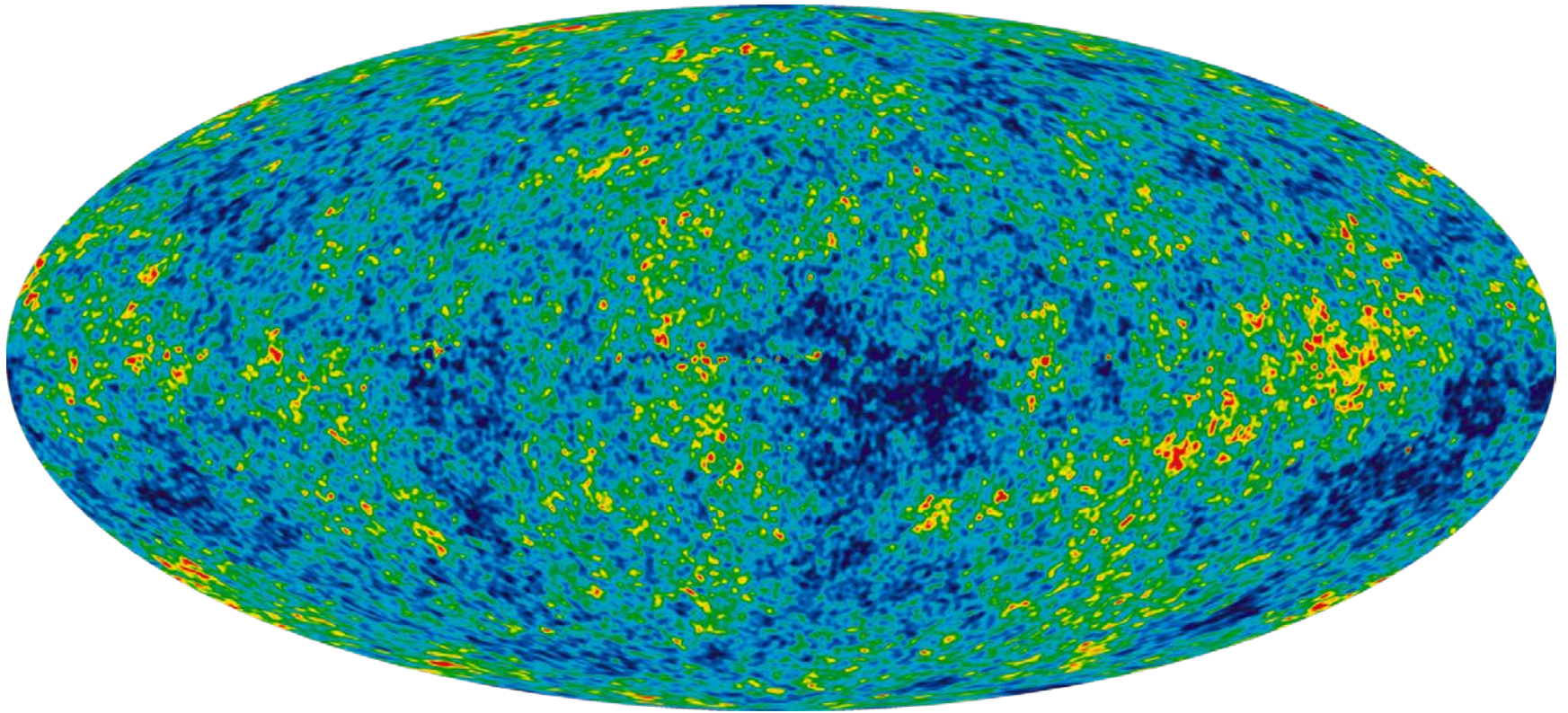
*eventually produce large scale structure of the Universe – stars, galaxies, galaxy clusters*

*have a logarithmic or power spectrum – amplitude increases by a few percent when the wavelength, increases several times*

$$\Phi \sim \lambda^{1-n_s}$$

$n_s$  --- spectral index

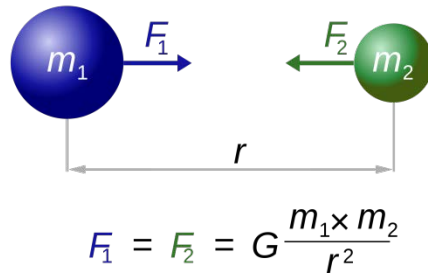




SKY <http://map.gsfc.nasa.gov/media/121238/index.html> The detailed, all-sky picture of the infant universe created from nine years of WMAP data. The image reveals 13.77 billion year old temperature fluctuations (shown as color differences) that correspond to the seeds that grew to become the galaxies. This image shows a temperature range of  $\pm 200$  microKelvin.  
Credit: NASA / WMAP Science Team WMAP # 121238 Image.

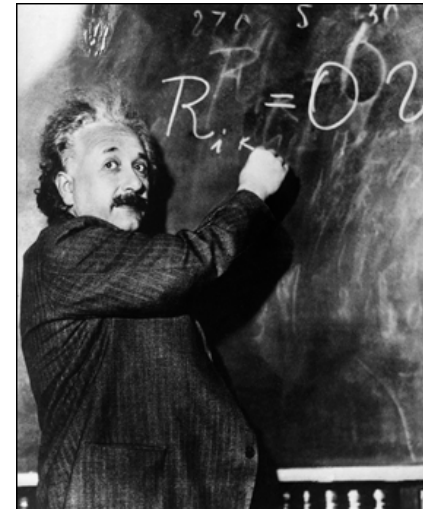
$$n_s = 0.968 \pm 0.006 \quad \text{Planck 2015}$$

# How comes that inflation expansion accelerates when gravity is an attraction force?

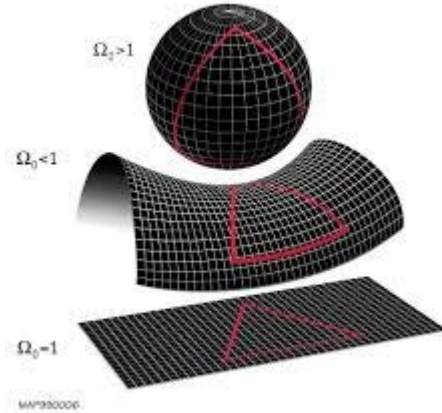


$$R_{\mu\nu} - \frac{1}{2}R g_{\mu\nu} + \Lambda g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

“Spacetime curvature” = “energy density and pressure”

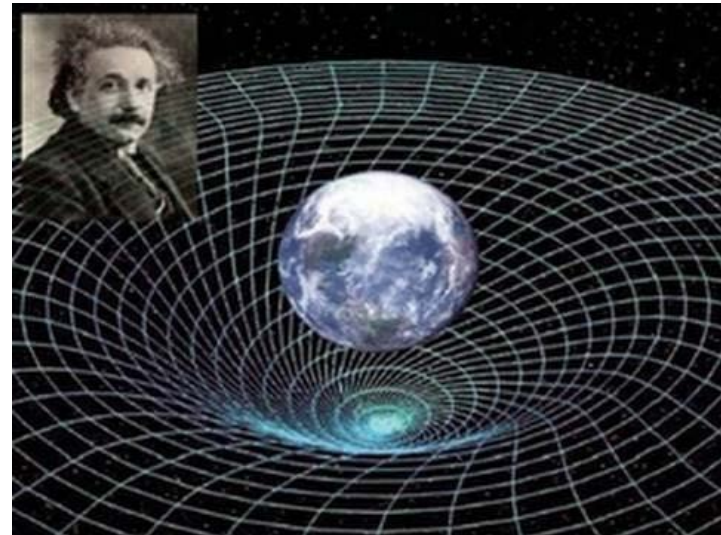


# What is spacetime curvature?



positive, negative and zero curvature of a 2-dimensional surface

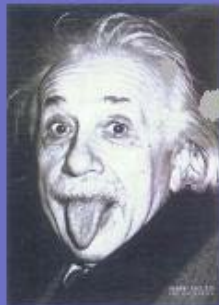
“Picture” of 4-dimensional curved spacetime:





What is  $\Lambda$  ?  $R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = \frac{8\pi G}{c^4} T_{\mu\nu} - \Lambda g_{\mu\nu}$

## The Cosmological Constant



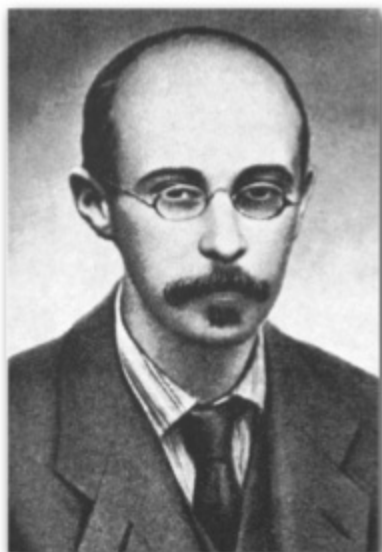
$$H^2 = \frac{8\pi G\rho}{3} - \frac{k}{a^2} + \frac{\Lambda}{3}$$

- ✦ Einstein introduced a **cosmological constant**  $\Lambda$  (vacuum energy) into GR to ensure a static universe.
- ✦ By tuning  $\Lambda$ , attractive gravity due to matter density can be balanced by the “repulsive” gravity, or negative pressure, of  $\Lambda$

Positive  $\Lambda$  -- repulsive force

# Alexander Friedman

1888 - 1925



“On the possibility of a world with constant negative curvature of space” (1924)

What happens if the Cosmological Constant ( $\Lambda$ ) has different values?

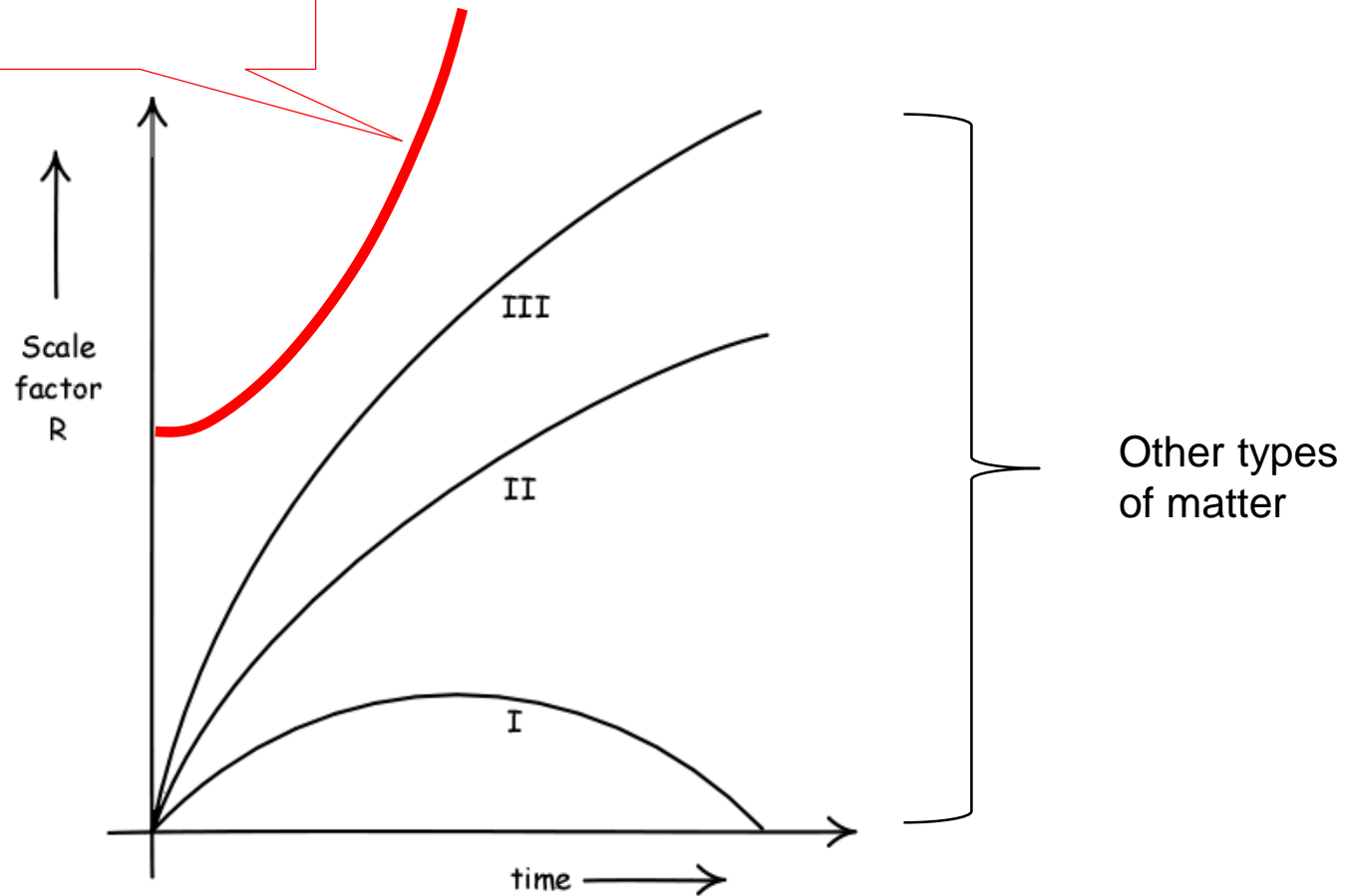
What happens if the universe is expanding?

$$R \sim e^{Ht}$$

$$H = \sqrt{\frac{\Lambda}{3}}$$

**Fast exponential growth of the size of the Universe**

**Hubble constant – the parameter of Hubble law in the *early* Universe**



- Einstein GR:**
- 1) Pressure also weighs
  - 2) pressure can be negative
  - 3) pressure can play the role of  $\Lambda$

**Scalar field with a potential  $V(\phi)$**

$$\varepsilon = \frac{\dot{\phi}^2}{2} + V(\phi)$$

$$p = \frac{\dot{\phi}^2}{2} - V(\phi)$$



**Slow roll approximation**  $\dot{\phi} \simeq 0$

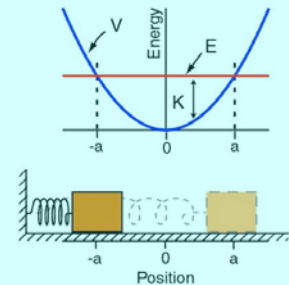
$$p = -\varepsilon, \quad \varepsilon = \frac{\Lambda}{8\pi G}$$

$$\Lambda = 8\pi G V(\phi) \simeq \text{const}$$

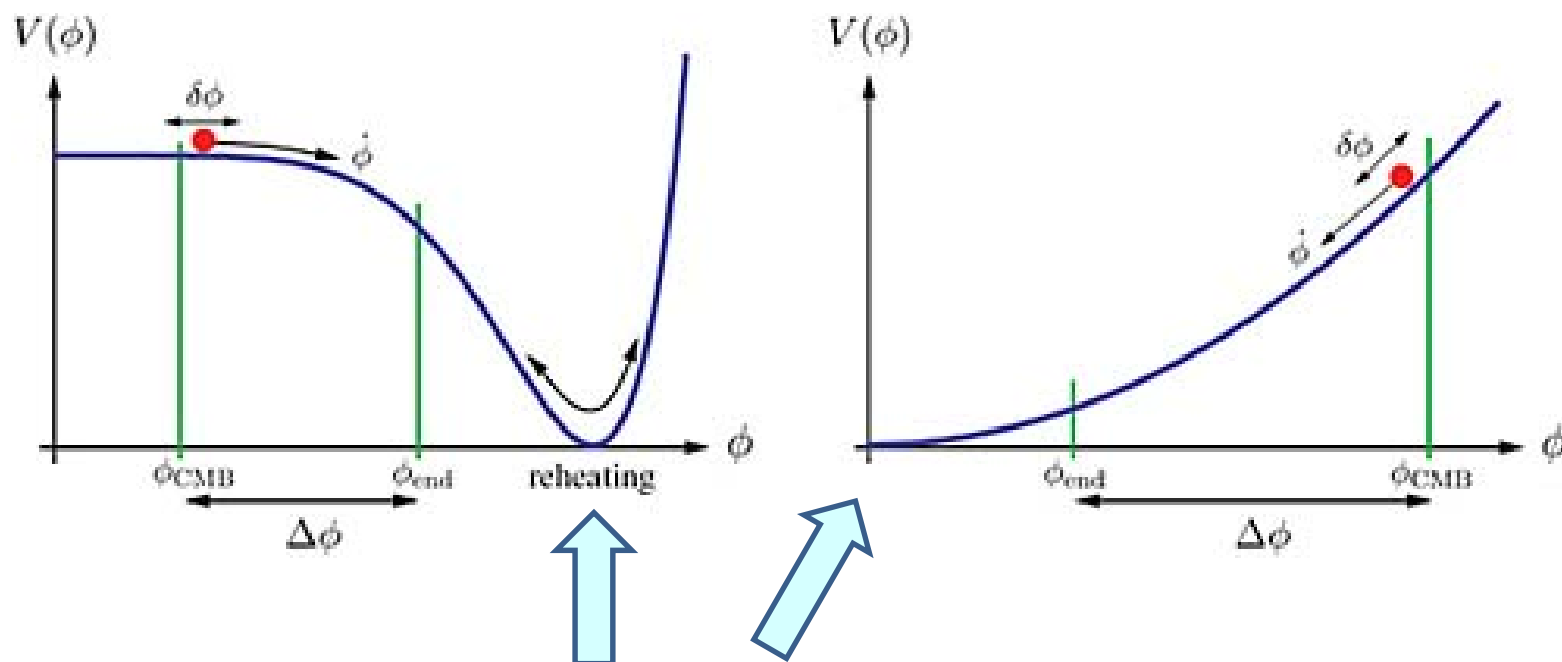
Mechanism of inflation --- slow roll of the  
**INFLATON** field  $\phi$  (A.Starobinsky, A.Linde)

Harmonic oscillator  
with quadratic potential

$$E = \frac{\dot{x}^2}{2} + \frac{\omega^2 x^2}{2}$$

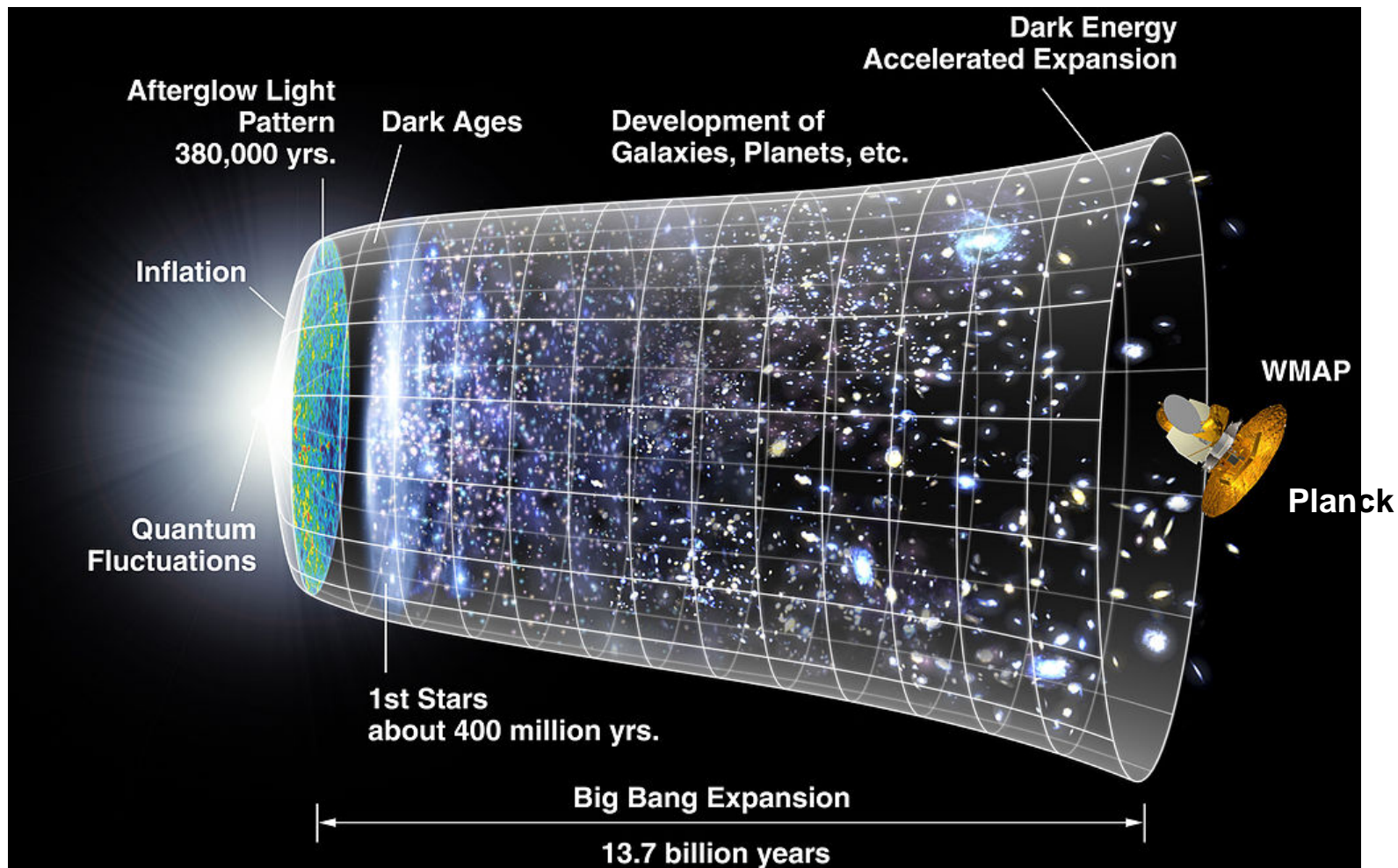


Shapes of potential -- types of inflation: “new” inflation, “chaotic” inflation, etc.



End of inflation, inflaton oscillations → reheating of the matter  
→ further standard model of Big Bang





[https://upload.wikimedia.org/wikipedia/commons/6/60/CMB\\_Timeline75.jpg](https://upload.wikimedia.org/wikipedia/commons/6/60/CMB_Timeline75.jpg)

By NASA/WMAP Science Team [Public domain], via Wikimedia Commons

# Candidate for the inflaton – Higgs boson

mass →	≈2.3 MeV/c <sup>2</sup>	≈1.275 GeV/c <sup>2</sup>	≈173.07 GeV/c <sup>2</sup>	0	≈126 GeV/c <sup>2</sup>
charge →	2/3	2/3	2/3	0	0
spin →	1/2	1/2	1/2	1	0
	<b>u</b> up	<b>c</b> charm	<b>t</b> top	<b>g</b> gluon	<b>H</b> Higgs boson
QUARKS	≈4.8 MeV/c <sup>2</sup>	≈95 MeV/c <sup>2</sup>	≈4.18 GeV/c <sup>2</sup>	0	
	-1/3	-1/3	-1/3	0	
	1/2	1/2	1/2	1	
	<b>d</b> down	<b>s</b> strange	<b>b</b> bottom	<b>γ</b> photon	
LEPTONS	0.511 MeV/c <sup>2</sup>	105.7 MeV/c <sup>2</sup>	1.777 GeV/c <sup>2</sup>	91.2 GeV/c <sup>2</sup>	
	-1	-1	-1	0	
	1/2	1/2	1/2	1	
	<b>e</b> electron	<b>μ</b> muon	<b>τ</b> tau	<b>Z</b> Z boson	
GAUGE BOSONS	<2.2 eV/c <sup>2</sup>	<0.17 MeV/c <sup>2</sup>	<15.5 MeV/c <sup>2</sup>	80.4 GeV/c <sup>2</sup>	
	0	0	0	±1	
	1/2	1/2	1/2	1	
	<b>ν<sub>e</sub></b> electron neutrino	<b>ν<sub>μ</sub></b> muon neutrino	<b>ν<sub>τ</sub></b> tau neutrino	<b>W</b> W boson	



[T.W.B.Kibble](#), [G.S.Guralnik](#),  
[C.R.Hagen](#), [F.Englert](#),  
[R.Brout](#) (+ P. Higgs) (1964)



F. Englert, P.Higgs  
Nobel Prize, 2013



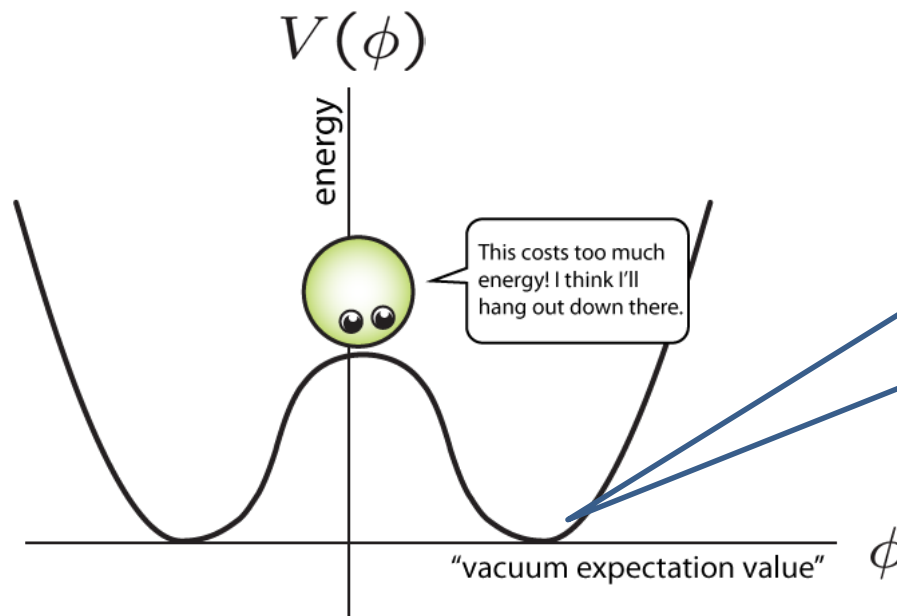
Large Hadron Collider  
CERN, Geneva  
July 2012

$M_{\text{Higgs}} \approx 125 \text{ GeV}$

# *The Higgs Boson*

The search for the God Particle

## Why “God particle”?

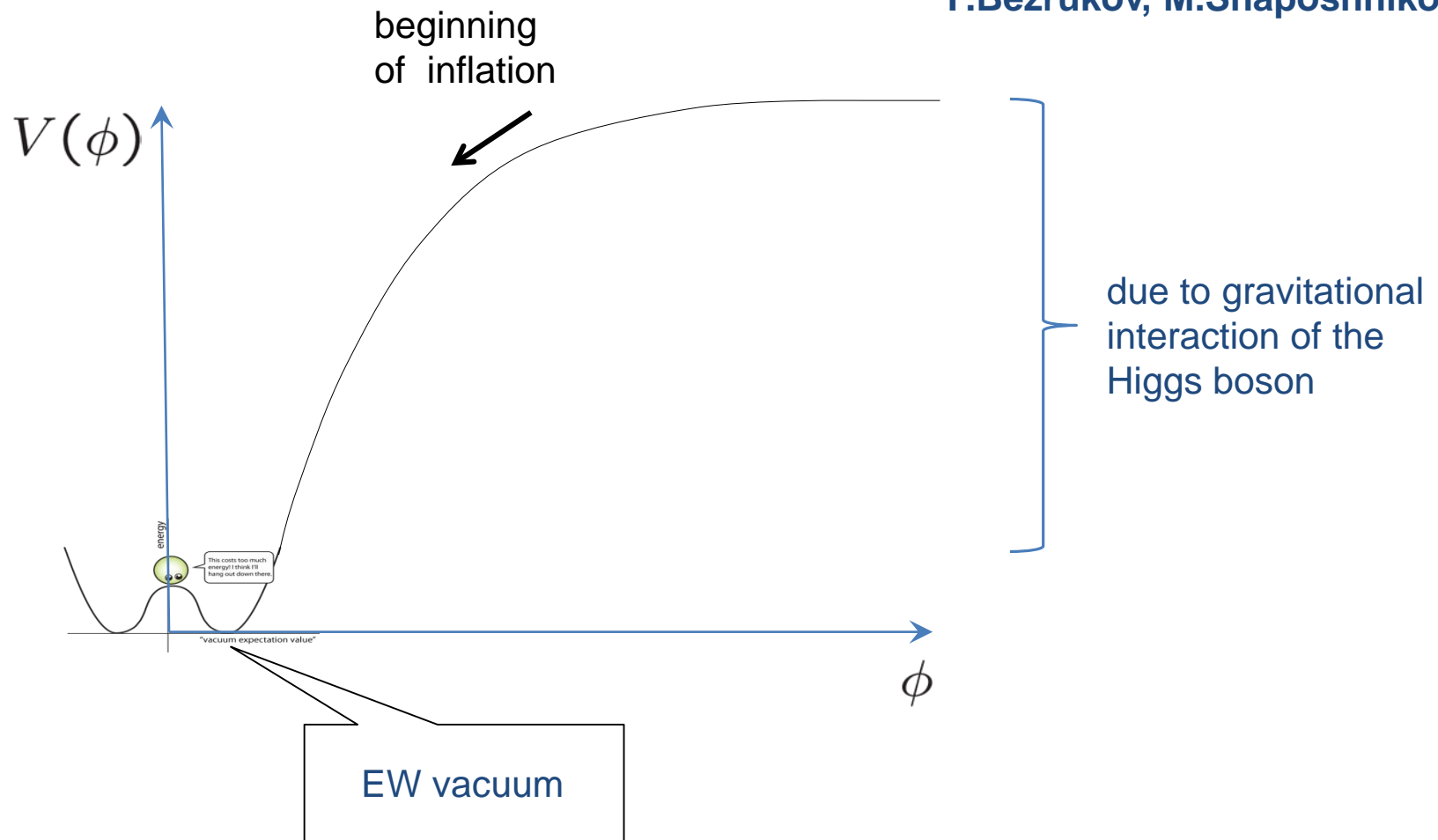


Higgs boson endows other particles with masses via the vacuum expectation value  $\hat{A}$  at the minimum of  $V(\hat{A})$  (electro-weak vacuum)

# Higgs inflation – Higgs boson coupled to gravity



F. Bezrukov, M. Shaposhnikov



$$n_s \simeq 0.96$$

**prediction of the theory**

$$M_{\text{Higgs}} \simeq 125 \text{ GeV}$$

## Unification of scales mediated by inflation:

Galactic  
scales

$$10\text{-}100 \text{ Mpc} \gg 10^{25} \text{ cm}$$

Electroweak  
scale

$$\gg 10^{-16} \text{ cm}$$



A.Kamenshchik, A.Starobinsky  
& A.Barvinsky 2008

A.Kamenshchik, C.Kiefer,  
A.Starobinsky, C.Steinwachs  
& A.Barvinsky. 2009

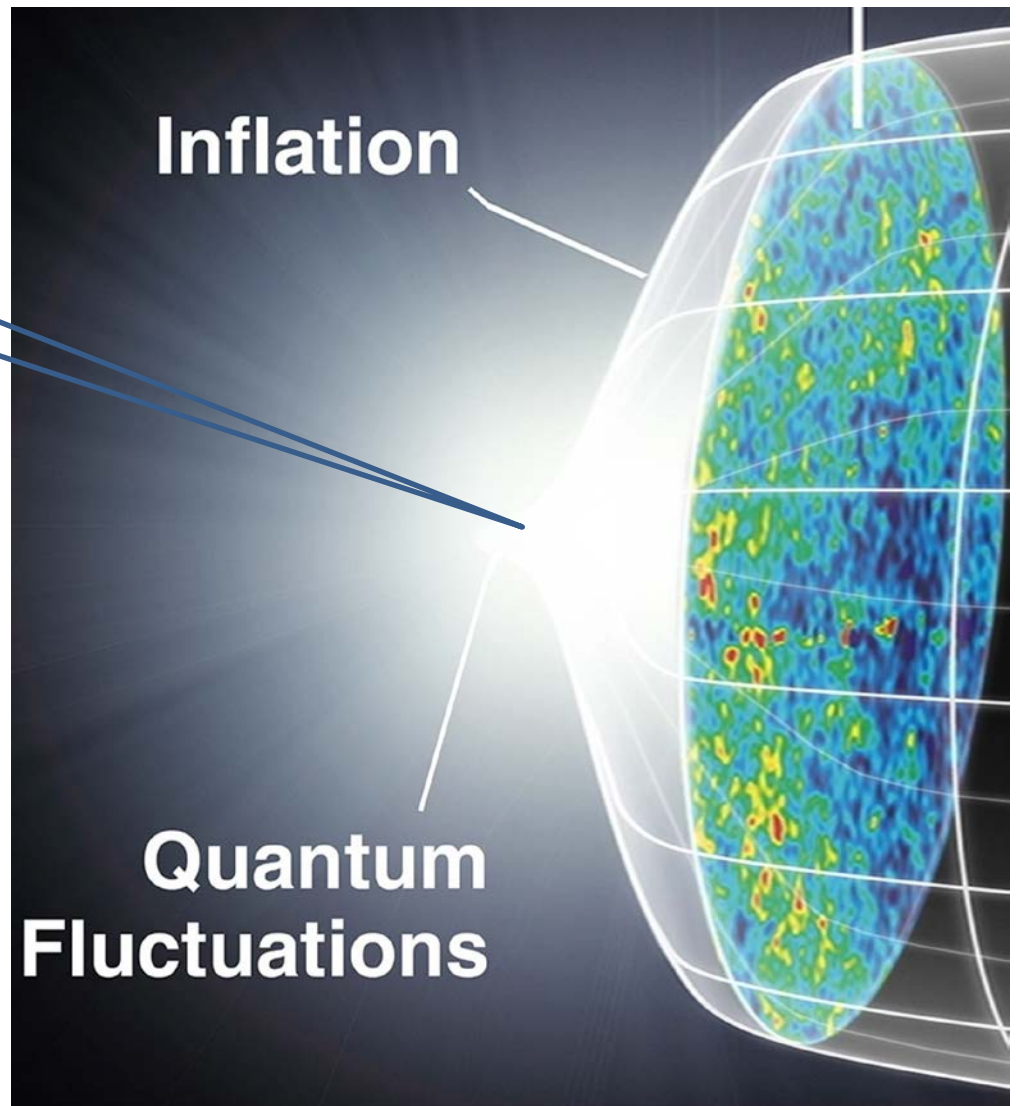
Bezrukov, Shaposhnikov, 2008



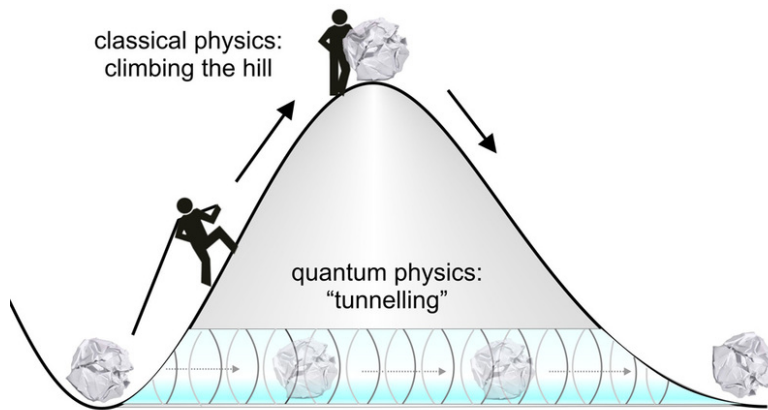
## What was at the beginning?

*The space and time had  
both one beginning . The  
space was made not in  
time but simultaneously  
with time.*

*Saint Augustin of Hippo*

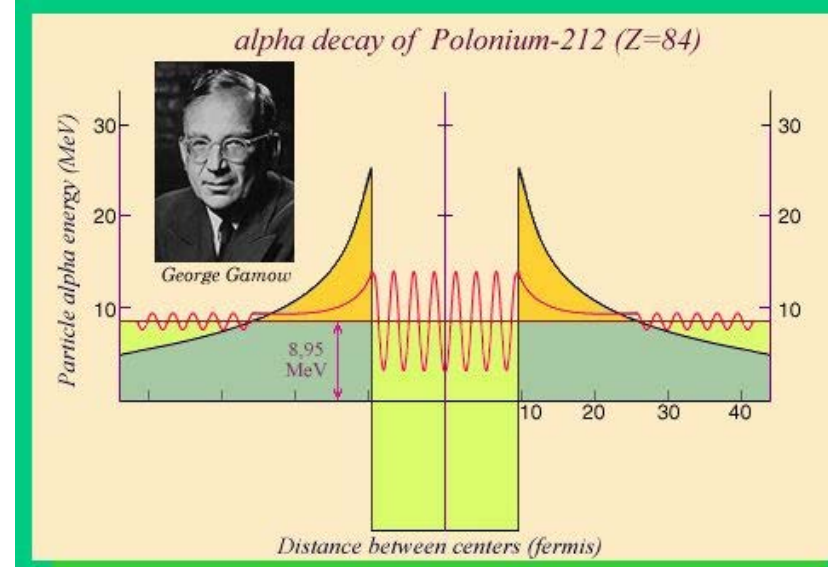
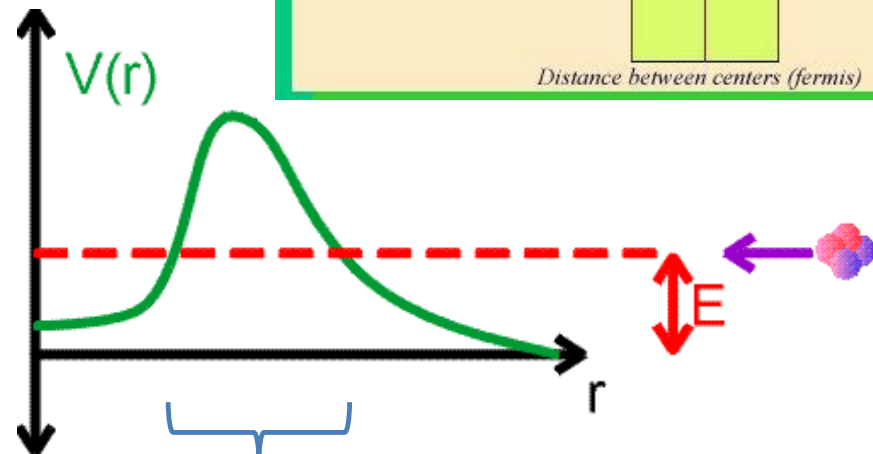


# Tunneling in QM:

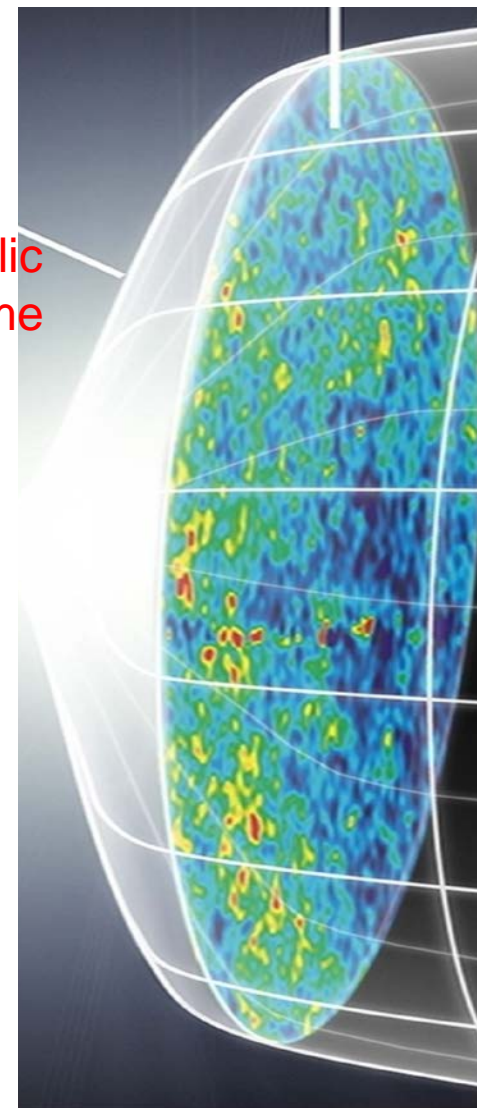
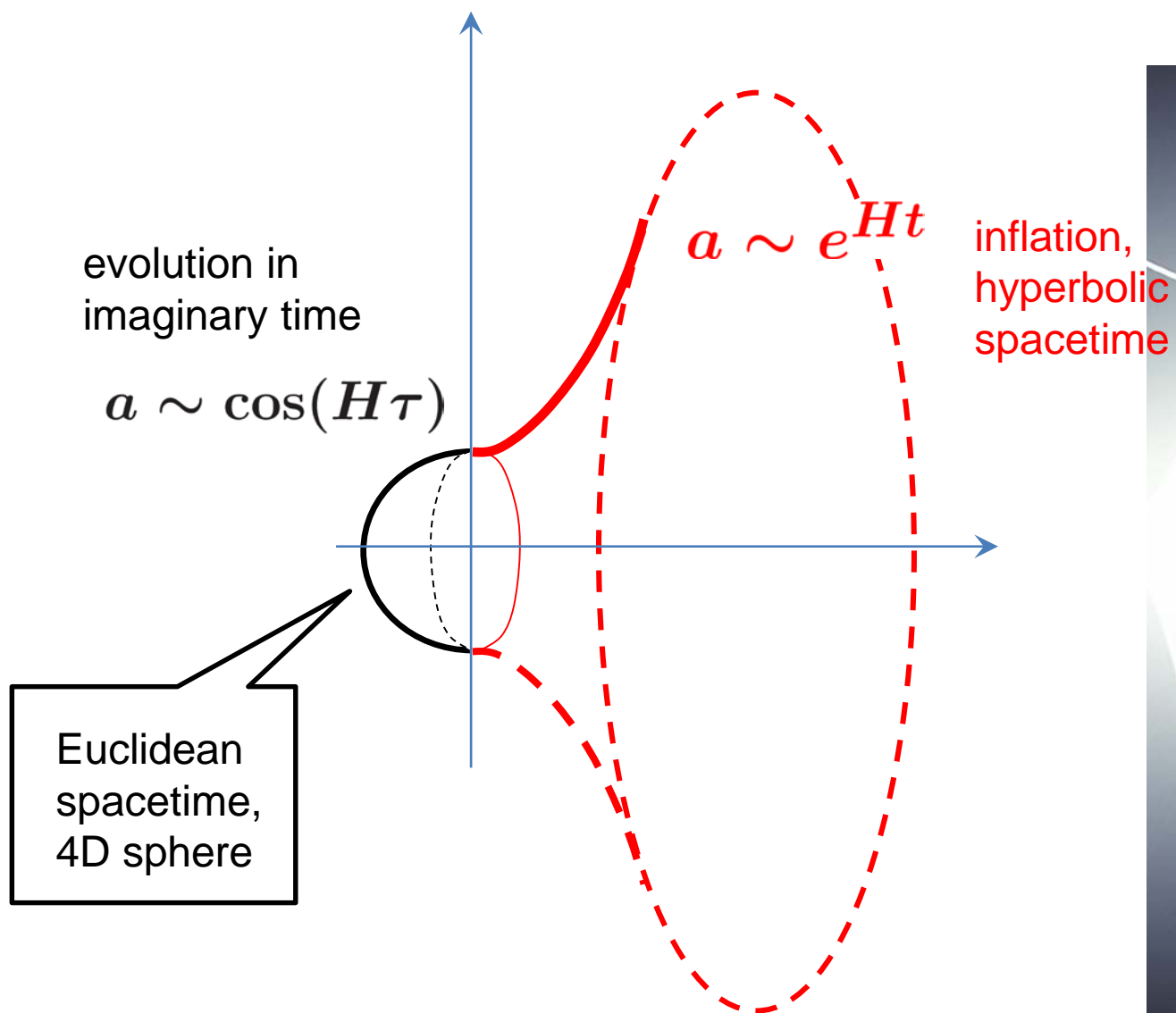


$$E = \frac{m\dot{r}^2}{2} + V(r)$$

$$\frac{dr}{dt} = \pm \sqrt{2 \frac{E - V(R)}{m}} = \text{imaginary}, \quad t = i\tau, \quad E - V < 0$$

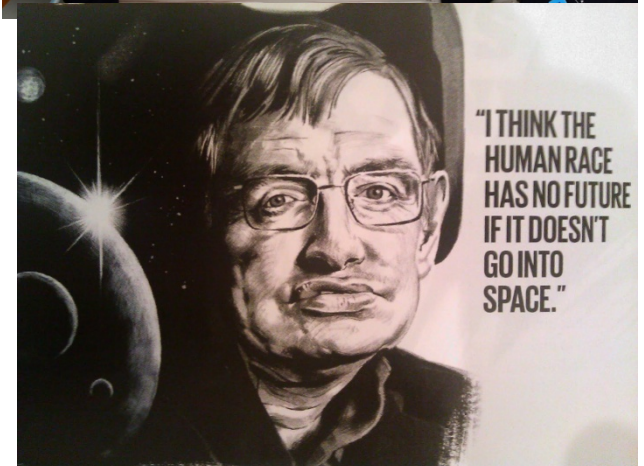
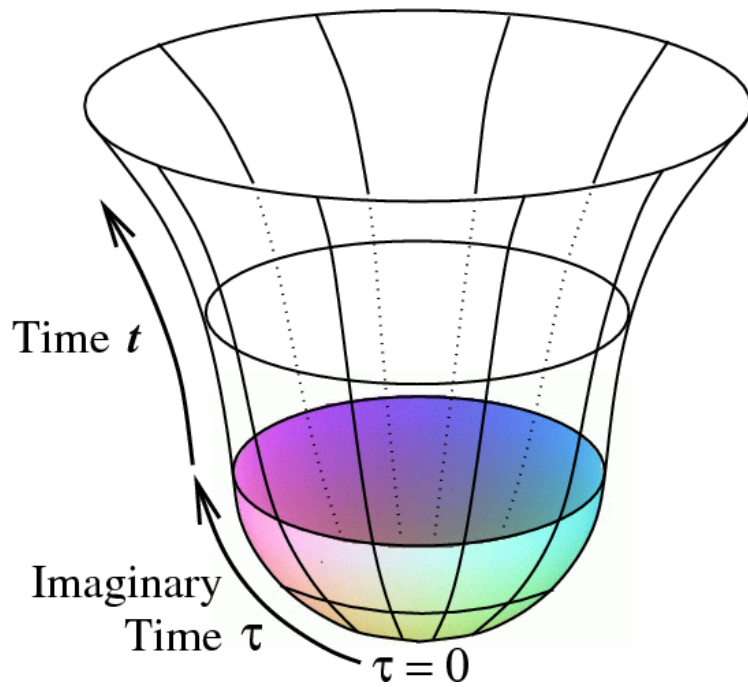
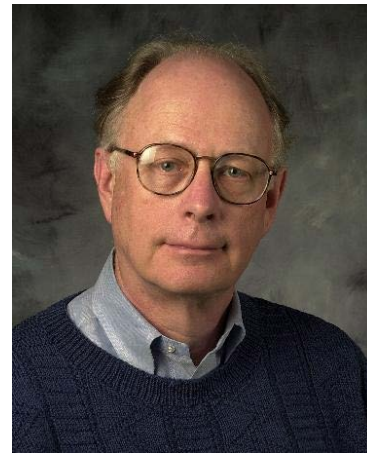


Underbarrier (classically forbidden) evolution in **imaginary** time  $\zeta$



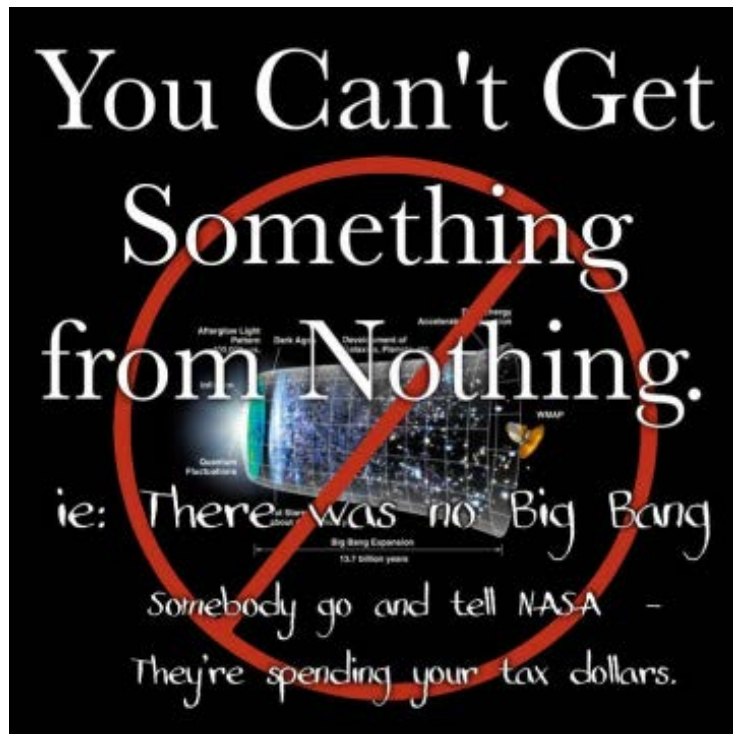
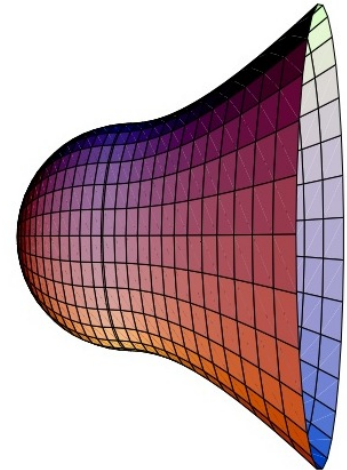


# Hartle-Hawking wavefunction of the Universe



# What might be wrong with the Hartle-Hawking wavefunction ?

Hartle-Hawking wavefunction is a **vacuum** state  
– the birth of the Universe from “**nothing**”



A vacuum state (as any other single state) is not unique – so why from “nothing” rather than from “Everything”?

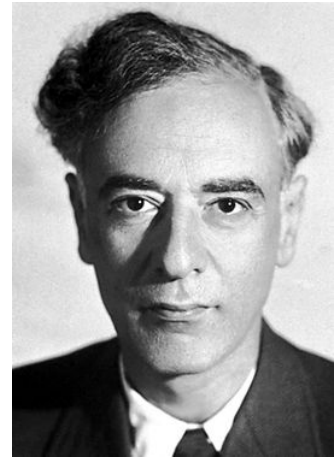
HH state does not generate good initial conditions for inflation – at the minimum of the inflaton potential rather than its maximum

**So what can be the alternative?**

This is a concept of transition from the wavefunction to the *density matrix* of the Universe:

$$\Psi \longrightarrow \hat{\rho}$$

# What is density matrix?



L.Landau

In classical theory **probabilities** sum up:  $p_1, p_2 \rightarrow p_1 + p_2$

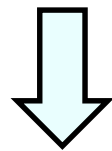
In quantum theory quantum **states (wavefunctions)** sum up:

$$|\Psi\rangle = |\Psi_1\rangle + |\Psi_2\rangle$$

$$\begin{aligned} \langle \Psi | \mathcal{O} | \Psi \rangle &= \langle \Psi_1 | \mathcal{O} | \Psi_1 \rangle + \langle \Psi_2 | \mathcal{O} | \Psi_2 \rangle \\ &\quad + \langle \Psi_1 | \mathcal{O} | \Psi_2 \rangle + \langle \Psi_2 | \mathcal{O} | \Psi_1 \rangle \end{aligned}$$

$\swarrow p_1$                        $\swarrow p_2$

cross terms – interference  
between red and blue



$$\Psi \rightarrow \hat{\rho}$$

$$\begin{aligned} \langle \Psi | \mathcal{O} | \Psi \rangle &= \langle \Psi_1 | \mathcal{O} | \Psi_1 \rangle + \langle \Psi_2 | \mathcal{O} | \Psi_2 \rangle \\ &\quad + \cancel{\langle \Psi_1 | \mathcal{O} | \Psi_2 \rangle} + \cancel{\langle \Psi_2 | \mathcal{O} | \Psi_1 \rangle} \end{aligned}$$

destruction of interference  
-- **decoherence**

## Rules of the game:

$$|\Psi\rangle = |\Psi_1\rangle + |\Psi_2\rangle \rightarrow \hat{\rho} = |\Psi_1\rangle\langle\Psi_1| + |\Psi_2\rangle\langle\Psi_2|$$

$$\langle\Psi|\mathcal{O}|\Psi\rangle \rightarrow \text{tr}(\mathcal{O}\hat{\rho}) \quad \text{trace operation}$$

$$\text{tr}(\mathcal{O}|\Psi_1\rangle\langle\Psi_1|) = \langle\Psi_1|\mathcal{O}|\Psi_1\rangle$$

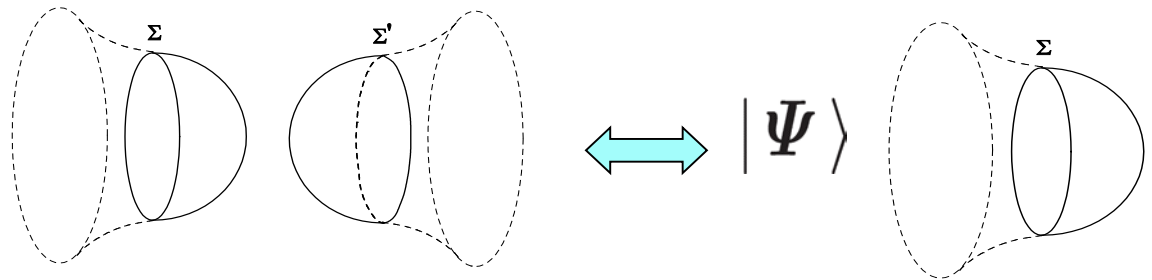
pure state  $|\Psi\rangle \Leftrightarrow \hat{\rho} = |\Psi\rangle\langle\Psi|$  *pure* state density matrix

$$\hat{\rho} = |\Psi_1\rangle\langle\Psi_1| + |\Psi_2\rangle\langle\Psi_2| + \dots$$

mixed state – *nonfactorizable*  
density matrix

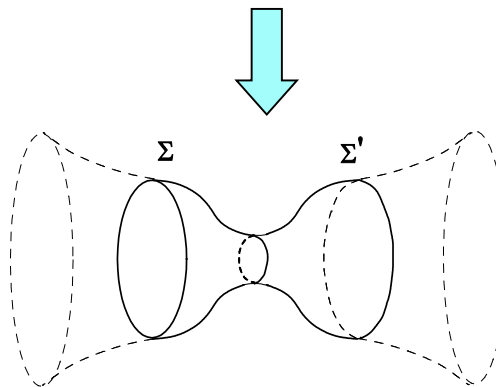
# From “vacuum” Hartle-Hawking state to the density matrix:

$$|\Psi\rangle\langle\Psi| = \hat{\rho}_{HH}$$



density matrix picture of a pure  
Hartle-Hawking state – vacuum state  
*of zero temperature*

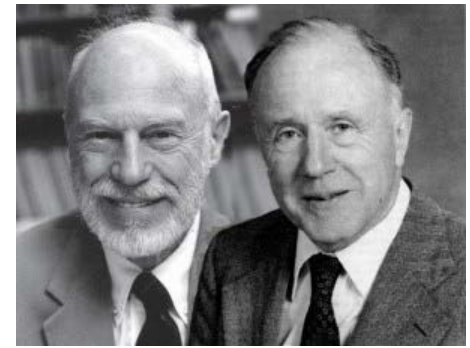
$\hat{\rho}_{\text{mixed}}$



density matrix picture of a mixed state

All possible physical states (wavefunctions) and density matrices in quantum gravity satisfy the **Wheeler-DeWitt equation** which is symbolically

$$\hat{H} |\Psi\rangle = 0$$



$$\hat{\rho} = \sum_{\text{all } |\Psi\rangle} |\Psi\rangle\langle\Psi|$$

A.Barvinsky., Phys. Rev. Lett. 99,  
071301 (2007)

Sum over “everything” that satisfies  
the Wheeler-DeWitt equation

Motivation: aesthetic (minimum set of assumptions – Occam razor)

An ultimate equipartition in the full set of states of the theory --- “**Sum over Everything**”.

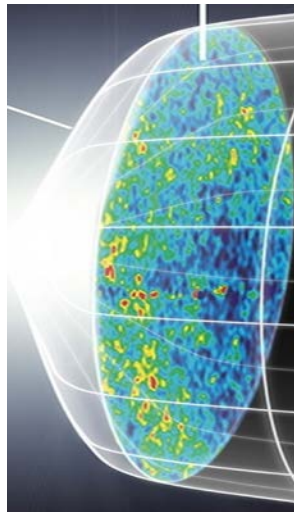
Creation of the Universe from **Everything** is conceptually more appealing than creation from **Nothing**, because the democracy of the microcanonical equipartition better fits the principle of the Occam razor than the selection of a concrete state.



# Properties of the “newly born” Universe

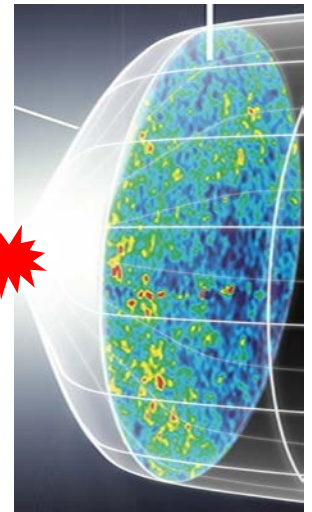
1) Initial thermal state with the primordial temperature  $T_{prim}$  of matter

## Standard inflation scenario **VS** Density matrix scenario



Vacuum,  
absolute zero  
temperature  
 $T=0$

Inflation, hot  
big-bang  
→ relic radiation  
 $T' 3000^{\circ}$



Thermal state,  
primordial  
temperature  $T_{prim}$

Inflation,  $T_{prim} \rightarrow 0$ ,  
hot big-bang  
→ relic radiation  
 $T' 3000^{\circ}$

# “SOME LIKE IT HOT” (SLIH) scenario



Known inflation paradigm retracted the BB concept by replacing it with the initial vacuum state.

“SOME LIKE IT HOT”  
-- it incorporates evolution.

**J**ournal of **C**osmology and **A**stroparticle **P**hysics **Not Big Bang**  
An IOP and SISSA journal **logical**

Cosmological landscape from nothing:  
some like it hot

A O Barvinsky<sup>1</sup> and A Yu Kamenshchik<sup>2,3</sup>

## 2) Restriction of the range of the effective cosmological constant below the Planck energy scale

$$\Lambda \leq \Lambda_{\max} \ll M_P^2 \sim (10^{18} \text{GeV})^2$$



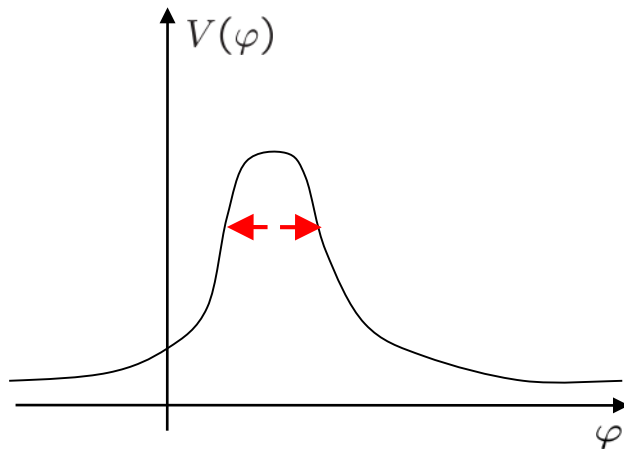
new quantum  
gravity scale

We do not have a consistent nonperturbative quantum gravity and have to use the powerful approximation method – semiclassical expansion – which is valid only below the Planck scale. This restriction of  $\Lambda$  justifies this method and its predictions!

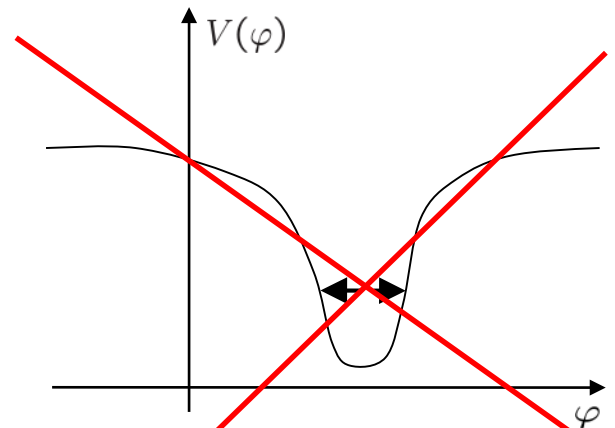
In string theory (“Theory of everything”) this property might be provoke the solution of the so called “landscape problem” (selection among  $10^{500}$  vacua).

**TOO  
VERBOSE**

### 3) Selection of inflaton potential *maxima* as initial conditions for inflation -- the major difficulty with the Hartle-Hawking state



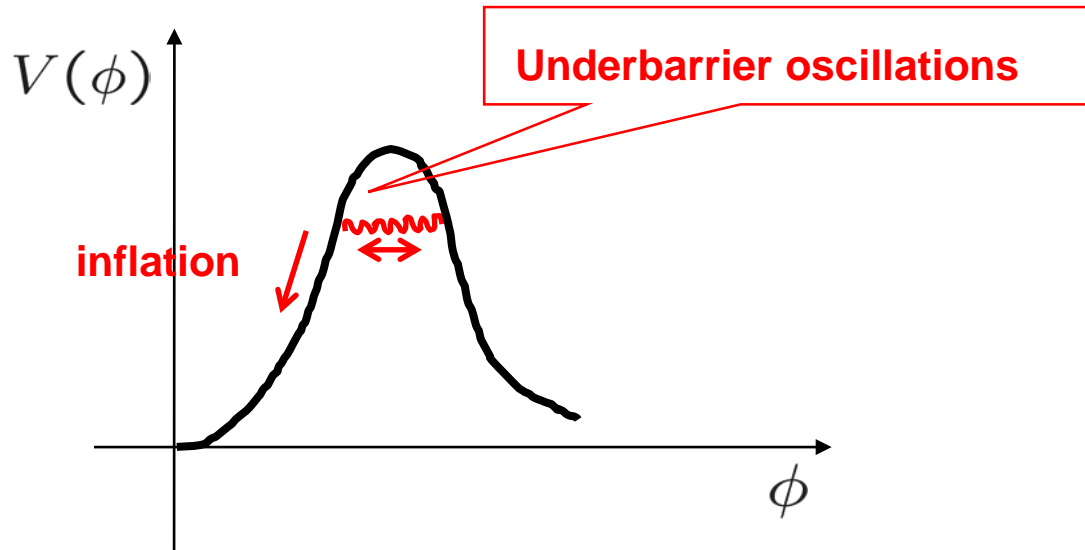
classically forbidden  
(underbarrier) oscillation



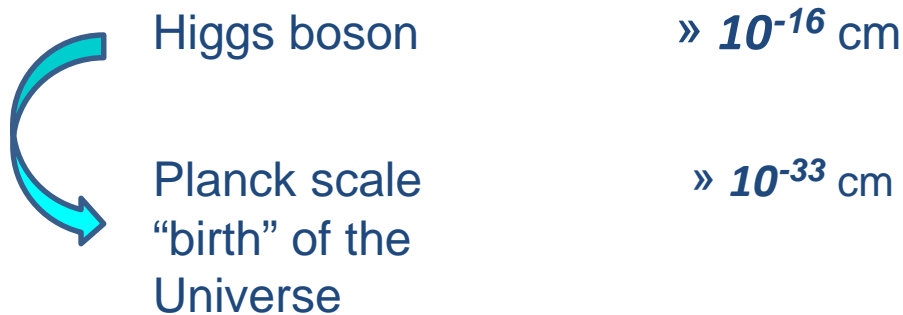
classically allowed (overbarrier)  
oscillation are ruled out

Initial conditions for a new type of hill-top inflation

## Picture of a new type hill-top inflation



Origin of the hill-top potential from quantum effects:  
remarkable match with the Higgs inflation model

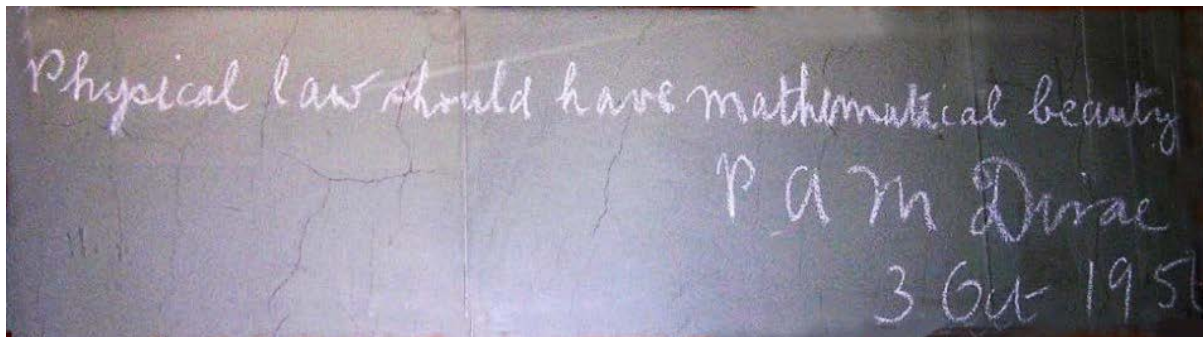


Our Universe embraces phenomena from quantum gravity –Planck  
-- scale  $10^{-33}$  cm to the Universe size of billions of light years

Unification of these phenomena is mediated by the transition from  
the quantum birth of the Universe at the Planck scale to the Higgs  
boson scale  $10^{-16}$  cm and further on to galactic scales

Our comprehension of this remarkable unity runs via the CMB  
observations -- gigantic microscope device provided to us by  
Nature in the form of inflation, the mechanism of enormous  
stretching of physical scales from  $10^{-16}$  cm to hundreds of Mpc

Remarkable beauty of this unification – absence of the redundant  
(in the spirit of Occam razor) culminates in the saying:



**TOO  
VERBOSE**